

# Handling non-Gaussian and nonlinear issues with variational methods

Elías Valur Hólm

ECMWF

May 11, 2010

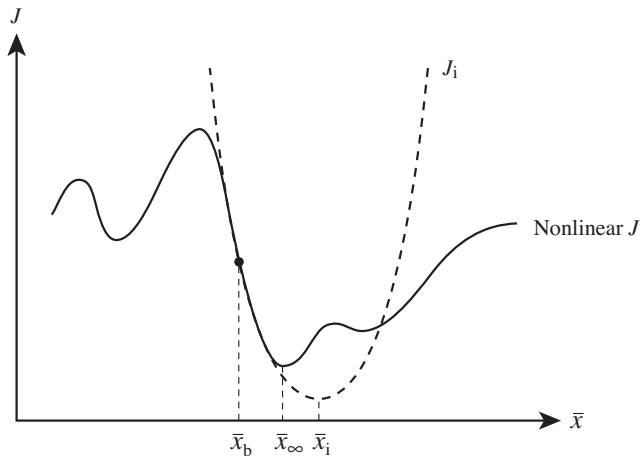
# Outline

- 1 Sources of nonlinearity
- 2 Model operator with tangent-linear model and physics
- 3 Observation operators
- 4 Observation error modelling
- 5 Background error modelling
- 6 Gaussianization of background/observation pdfs

# Sources of nonlinearity in variational data assimilation

- The full **nonlinear problem** is solved as a **sequence of linear problems** (outer loops), each of which is made up of several iteration of the inner loops using tangent linear models **M** and observation operators **H** at reduced resolution.
- In each iteration the costfunction is linearized around a previous estimate  $\mathbf{x}_i$ , which gives a quadratic costfunction  $J_i$ .
- For weakly nonlinear systems this procedure works well and is widely used.
- The convergence depends on:
  - ▶ the accuracy of the first estimate of  $\mathbf{x}$  (the background  $\mathbf{x}_b$ )
  - ▶ and the nonlinearity of  $J$ .
- There are Monte Carlo assimilation algorithms which deal directly with the nonlinearity. However, often this is not trivial for large dimensional problems (e. g. large number of particles/members needed for capturing nonlinearities with particle filters/ensemble Kalman filters)

# The inner/outer loop solution of the nonlinear problem



# How to handle model nonlinearities?

- The error in the tangent linear model perturbation at reduced resolution is estimated by comparison with differences of nonlinear integrations,

$$\mathbf{M}\delta\mathbf{x}_0 - (\mathcal{M}(\mathbf{x} + \delta\mathbf{x}_0) - \mathcal{M}(\mathbf{x}))$$

- Factors affecting this difference are:
  - ▶ **Number of outer loops** - each relinearization increases the chance of capturing nonlinearities (presently 3 at ECMWF).
  - ▶ **Resolution of the inner loops** - ideally same as outer loops, but needs to be reduced (at ECMWF factor 3 balances accuracy/efficiency).
  - ▶ **Length of the assimilation window** - longer windows decrease the accuracy of the TL approximation (presently 12-24 h at ECMWF).
  - ▶ **Model dynamics and physics** - chaotic transitions and on/off processes decrease linearity (at ECMWF we gradually enhance linearized physics).
  - ▶ **Size of the analysis increments** - smaller increments with better observations/analysis.

# Effect of model variable choice on nonlinearities

- Some model variables are more linear than others, e. g.:
  - ▶ Condensation/evaporation rapidly converts between humidity and condensed water, but the sum is more constant in time,  
 $(q, q_l, q_i) \longrightarrow q_t = q + q_l + q_i$ .
  - ▶ Stratospheric chemistry rapidly converts between  $NO$  and  $NO_2$  during the night, but the sum is more constant,  
 $(NO, NO_2) \longrightarrow NO_x = NO + NO_2$ .
- We can keep the original model variables in the nonlinear model  $\mathcal{M}$  and only use the combined variables in the TL model  $\mathbf{M}$  to make the linear approximation better.
- We can also only use the combined variables in the control variable with the aim to get more Gaussian statistics. Or use a different combination for the control variable, e. g. statistics for  $\delta q_u = \delta q - \alpha \delta T$  taking into account correlations and balances!
- Lesson is that the correct answer for any model depends on your own research to find the variable giving best linearity and accuracy.

# How to handle observation operator nonlinearities?

- The error in the tangent linear observation operator can be estimated by comparing with the difference between two nonlinear observation operator calculations,

$$\mathbf{H}\delta\mathbf{x}_i - (\mathcal{H}(\mathbf{x}_i + \delta\mathbf{x}_i) - \mathcal{H}(\mathbf{x}_i))$$

- Factors affecting this difference are:
  - ▶ The linearity of the observation operator.
  - ▶ Resolution of the inner loops.
  - ▶ The length of the assimilation window.
  - ▶ The model physics.
  - ▶ The size of the analysis increments - smaller increments will reduce the agitation of nonlinearities above.

# The linearity of the observation operator

- Many operators are linear or nearly linear, like those just requiring interpolation of model variables, e. g. airplane temperature.
- Some operators include **physical on/off processes**, like cloud and rain observations, which require a careful formulation of linear physics: microwave radiances and radar.
- Infrared radiances are sensitive to the presence of clouds and sea ice, **abrupt changes** to radiance depending on whether it is clear/cloudy with/without ice.

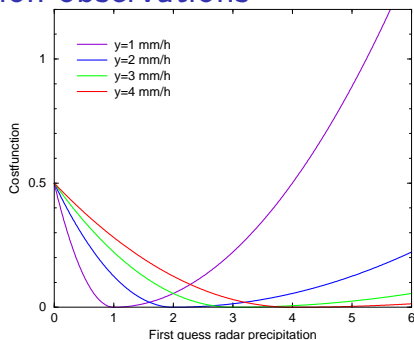
# Sources of nonlinear observation errors

- Observation errors are often accurately modelled by Gaussian distributions, e. g. radiosonde temperature errors.
- Non-Gaussian observation error distributions have **different sources**:
  - ▶ The observations include **bad observations** affected by systematic errors, causing skewed or even multimodal distributions - this is looked at in the lecture on **quality control**, e. g. Huber norm for robust estimation.
  - ▶ The observations contain different **populations with different error characteristics**, e. g. cloudy and non-cloudy infrared radiances, again possibly causing skewed/multimodal distributions.
  - ▶ The observations are not direct, but **retrievals**, with biases and ambiguities of the retrievals causing skewed/multimodal distributions, e. g. satellite winds from cloud motion vectors (height assignment not unique) and scatterometer winds, which have a multimodal mapping from the measured backscatter to winds.
  - ▶ The observations are of **bounded quantities**, causing skewed distributions, e. g. rain observations, humidity and some chemical tracer observations.

# How to model non-Gaussian observation errors?

- For **multi-modal distributions**, the first step is to identify the characteristics of the different sub-distributions and model the observation errors separately for each class of observations, e. g. cloudy vs. non-cloudy IR radiances.
- More generally, the **variance and bias of the observations may vary continuously as a function of some stratifying variable**, e. g. the **amplitude of the observation itself**. An example is rain observations.
- The **cause of skewed distributions also needs to be identified**. If caused by a superposition of different populations, the same techniques apply as for the multimodal case.
- Skewed distributions caused by **boundedness of the observations are modelled with asymmetric observation error distributions**, e. g. for rain observations there is larger chance for the error to be on the high than the low side when the observed value is low.

# Radar precipitation observations



- Example of **asymmetric costfunction** for precipitation (Koizumi, Ishikawa and Tsuyuki, JMA):

$$J_{rain} = \frac{1}{2r^2}(y - (\mathcal{H}_{rain}(\mathbf{x}_b)))^2$$

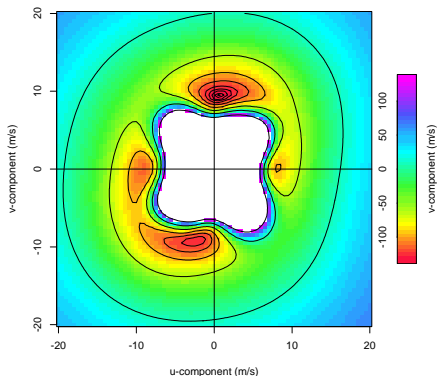
where  $r = r_c$  if  $\mathcal{H}_{rain}(\mathbf{x}_b) \leq y$ ,  $r = 3r_c$  if  $\mathcal{H}_{rain}(\mathbf{x}_b) > y$ , with  $r_c = \max(y, 1\text{mm/h})$ .

# How to handle non-Gaussian observation errors in the assimilation?

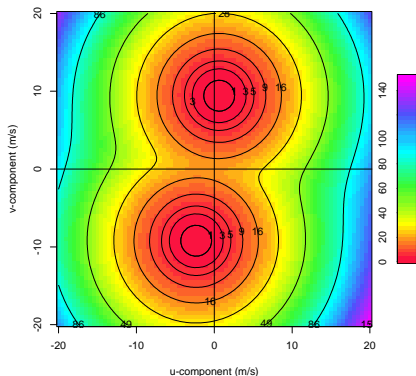
- Possibilities depend on the **minimization algorithm** used.
- For the **conjugate gradient** method, which requires a continuous Hessian and thus strictly quadratic costfunction, we must re-linearize  $J_o$  at each outer loop level.
- For the **quasi-Newton method**, like M1QN3 from INRIA, which can handle weak nonlinearities,  $J_o$  is either re-linearized w. r. t. the **observation error at outer loop level as above, or the nonlinear observation error is used in the inner loops**. This works especially well if the observation error model maintains a convex costfunction, which does not create any additional minima.
- Useful **probability distributions giving convex costfunction** include:
  - ▶ **Gaussians with different variances joined at zero**: Good for skewed distributions.
  - ▶ **Huber distribution**: A Gaussian in the middle is joined to exponential distributions to the left and to the right. Good for skewed and fat tail distributions with more frequent outliers than a Gaussian.

# Scatterometer wind observation costfunction

Maximum-Likelihood Estimator



Dual winds functional



- The costfunction for wind is multimodal (left), but can be simplified by combining the most probable solutions (right). (Courtesy Mark Leitner)

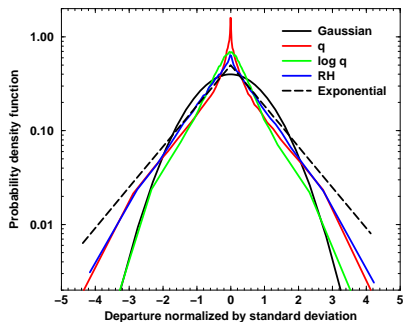
# How to handle non-Gaussian background errors?

- The **sources of nonlinearities coming from the model integration** producing a background state were discussed above.
- If the background errors become non-Gaussian, we need to treat these as **perturbations on top of the Gaussian framework**, if we are to stay with efficient algorithms which assume quadratic costfunctions. Generally works well in combination with the techniques discussed above to handle nonlinearities at outer loop level (re-linearization etc).
- The main problems arise when:
  - ▶ Variables **change very abruptly**, e. g. due to on/off physical processes: cloud variables.
  - ▶ Sharp **gradients in horizontal/vertical** cause anisotropies: boundary layer top, fronts.
  - ▶ Variables contain **scales spanning several orders of magnitude**: upper tropospheric humidity spans over 4 orders of magnitude.
  - ▶ Variables are **physically bounded within limits**: humidity, clouds, tracers.

# Anisotropic and flow-dependent horizontal and vertical correlations

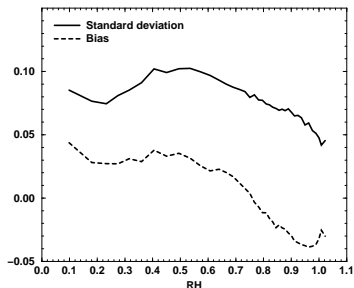
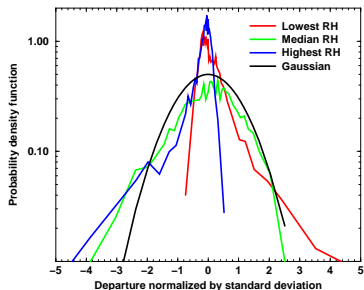
- Anisotropies can be introduced in several ways:
  - ▶ **Diffusion operator, digital filter and gridpoint** formulations of the background error all allow 3D anisotropy. Need specification of coefficients!
  - ▶ **Coordinate transforms** in both horizontal and vertical: geostrophic coordinate transform, boundary layer top and tropopause following vertical coordinates.
- The specification of coefficients needed to produce anisotropy can be made flow dependent (based on last outer loop trajectory).
- In general a **statistical regression of error correlations as function of variables/gradients** is needed, preferably guided by physical insights/principles.
- A natural extension is to include information from all ensemble members in an ensemble assimilation system.

# Symptoms of inhomogeneous background error variances



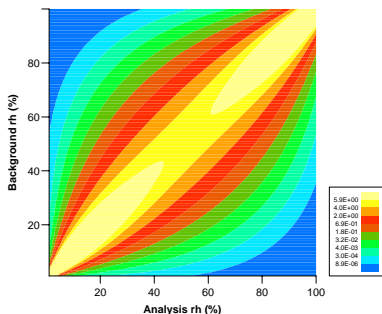
- Global humidity forecast difference histograms at  $850hPa$  model level show **peak at zero, where all the small errors gather** - typical of an inhomogeneous situation.
- Using different variables does not help much.

# Symptoms of bounded background error variances



- Relative humidity histograms as above for different bins of background relative humidity show asymmetry at boundaries, accompanied by biases.

# The joint pdf $P^b(rh^1, rh^2)$

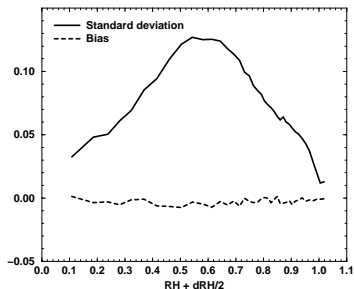
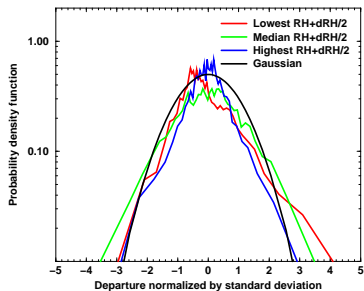


- The joint pdf (from forecast differences) is asymmetric w. r. t.  $rh^1$ , but symmetric w. r. t.

$$\frac{1}{2}(rh^1 + rh^2) = rh^1 + \frac{1}{2}\delta rh$$

- We have labelled the two forecasts analysis and background, but should be fc1/fc2.

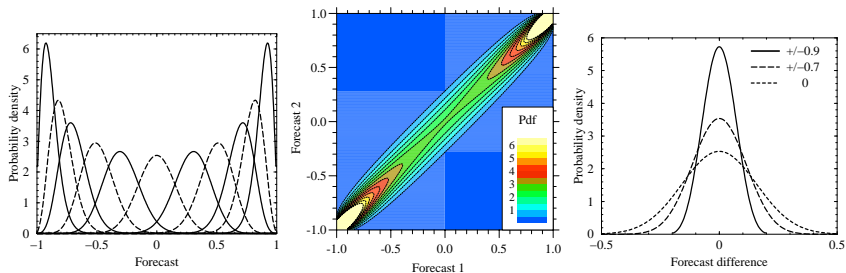
# Normalized and symmetrized humidity statistics



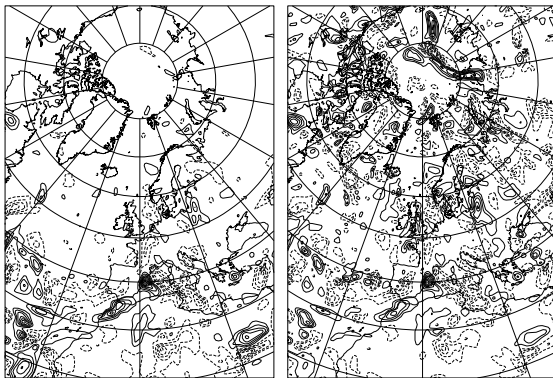
- The symmetric pdf  $P(\delta rh | rh^b + \frac{1}{2}\delta rh)$  can be modelled by a normal distribution
- The variance changes with  $rh^b + \frac{1}{2}\delta rh$ , and the bias is zero. A control variable with approximately a normal distribution  $\mathcal{N}(0, 1)$  is obtained by nonlinear normalization

$$\tilde{\delta rh} = \frac{\delta rh}{\sigma(rh^b + \frac{1}{2}\delta rh)}$$

# Example: Background error for a harmonic oscillator amplitude



## Effect of normalization on homogeneity



- Forecast differences (ca. 850 hPa) for specific humidity  $\delta q$  (left) and normalized relative humidity  $\widetilde{\delta rh}$  (right).
- The normalized relative humidity gives a more homogeneous field for the following normalization in spectral or wavelet space.

# Implementing nonlinear humidity background errors

- The background error costfunction  $J_b$  is now nonlinear

$$J_b = f^T(\delta rh) B^{-1} f(\delta rh)$$

$$f(\delta rh) = \frac{\delta rh}{\sigma(rh^b + \frac{1}{2}\delta rh)}$$

- Implementation requires linear inner loops with nonlinearities in outer loops

- ▶ For outer loop  $i$ , inner loops use  $\widetilde{\delta rh}^i = \delta rh^i / \sigma(\frac{1}{2}(rh^b + rh^{i-1}))$ .
- ▶ In outer loops  $\widetilde{\delta rh}$  is given. Solve  $\delta rh$  from the nonlinear equation

$$\frac{\delta rh}{\sigma(rh^b + \frac{1}{2}\delta rh)} - \widetilde{\delta rh} = 0$$

# Direct 'gaussianization' of background/observation pdfs

- Nonlinear error in both background and observation costfunctions can be transformed analytically to obtain a more Gaussian costfunction, e. g. logarithmic transform of the control variable. An analytic transform like this may not fit the statistics.
- More generally, a numerical transform based on forecast difference statistics can be derived, based on the pdf  $P(\delta\mathbf{x})$ , or even better the conditional pdf  $P(\delta\mathbf{x}|\Phi)$ , where  $\Phi$  is a suitable stratifying variable which separates large and small variances for fields containing different scales (like humidity or atmospheric tracers).
- For a given  $\Phi$  we just need to find a transformation  $f(\delta\mathbf{x}, \Phi)$  of the  $\delta\mathbf{x}$  axis so that the probability that  $\eta \leq \delta\mathbf{x}$  equals the probability that  $\xi \leq f(\delta\mathbf{x}, \Phi)$  for a normal gaussian distribution

# The 'gaussianization' transform using cumulative pdfs

- The cumulative pdf  $\Pi$  is related to the gaussian cumulative pdf  $\Pi_G$  by

$$\Pi(\delta\mathbf{x}|\Phi) = \int_{-\infty}^{\delta\mathbf{x}} P(\eta|\Phi)d\eta = \int_{-\infty}^{f(\delta\mathbf{x},\Phi)} \frac{1}{\sqrt{2\pi}} e^{-\xi^2/2} d\xi = \Pi_G(f(\delta\mathbf{x},\Phi))$$

- Inverting the gaussian cumulative distribution then gives

$$f(\delta\mathbf{x},\Phi) = \Pi_G^{-1}(\Pi(\delta\mathbf{x}|\Phi))$$

- This will always work, but we may end up with a bias term which does not fit into the standard analysis framework - the costfunction needs to be modified to account for the bias.
- However, we could also apply this transform as a finishing step after the symmetrizing transform and in that way avoid the bias term.